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HEIDELBERG
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SEIT 1386

New families of scale separated vacua

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Based on [2309.00043](#) [R. Carrasco, TC, F. Marchesano and D. Prieto] and review
[2311.12105](#) [TC]

1. Scale separation: Definition and current status
2. New families: Motivations
3. New families: More details

Scale separation: Definition and current status

Hiding the extra dimensions

10d spacetime =



4d world

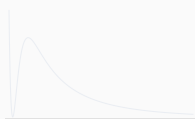
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Compact X^6

- Cosmological constant Λ

$$\Rightarrow R_{(\Lambda)\text{dS}} M_{\text{p}} \sim |\Lambda|^{-1/2}$$



- Kaluza-Klein scale $M_{\text{KK}} \sim \frac{M_{\text{s}}}{\text{Vol}_{X^6}^{1/6}}$ (and winding scale)

Scale separation: $R_{(\Lambda)\text{dS}} M_{\text{KK}} \gg 1$

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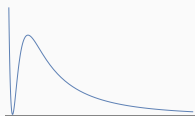
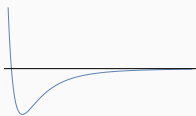
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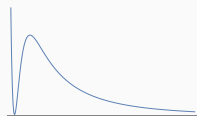
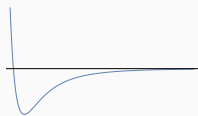
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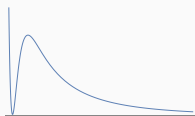
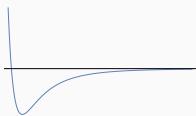
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- $X_6 = T^6 / (\mathbb{Z}_3 \times \mathbb{Z}_3) \implies h^{2,1} = 0$ & let us forget twisted moduli

$$t^a \propto \frac{\sqrt{|\hat{e}_1 \hat{e}_2 \hat{e}_3|}}{|\hat{e}_a| \sqrt{m}} \quad \text{with} \quad \hat{e}_a \equiv e_a - \frac{1}{2} \frac{\mathcal{K}_{abc} m^b m^c}{m}$$

Tadpole: $dG_2 = G_0 H + Q_{06} \delta_{06}$

\implies Flux quanta e_a, m^a unconstrained

$$\hat{e}_a \sim n$$

$$t^a \sim n^{\frac{1}{2}}, \quad \text{Vol}_{X_6} \sim n^{\frac{3}{2}}, \quad e^\phi \propto (\mathcal{K}_{abc} t^a t^b t^c)^{-\frac{1}{2}} \sim n^{-\frac{3}{4}}, \quad e^{\phi_4} \sim n^{-\frac{3}{2}}$$

Good control: $n \rightarrow \infty \implies$ large volume/weak coupling

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$$\Lambda \propto -3e^K |W|^2 \sim n^{-\frac{9}{2}}$$

$$R_{\text{AdS}} M_{\text{KK}} \sim n^{\frac{1}{2}}, \quad \text{and} \quad R_{\text{AdS}} M_{\text{W}} \sim n$$

Susy AdS:

- Moduli stabilised
- Controlled regime
- Scale separation
- Pert. stable

Generalisable:

- Metric fluxes [Camara, Font, Ibáñez '05]
- [Derendinger, Kounnas, Petropoulos, Zwirner '05] [Villadoro, Zwirner '05]
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- [Marchesano, Quirant '19]
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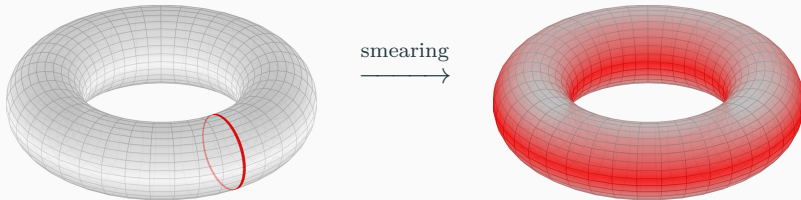
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10d uplift of DGKT cannot be CY

O6-planes are **diluted** in compact space [Acharya, Benini, Valandro '07]

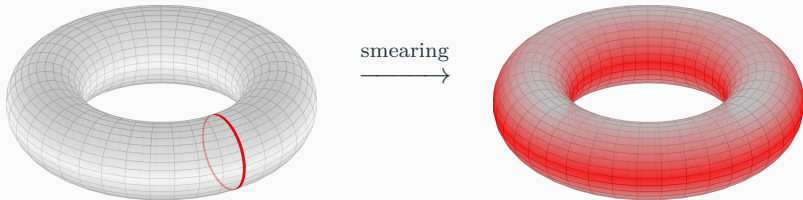


$$\delta_{O6}(\vec{y}) \propto \sum_{\vec{k}} e^{i\vec{k}\cdot\vec{y}} = 1 + \sum_{\vec{k}\neq\vec{0}} e^{i\vec{k}\cdot\vec{y}} \longrightarrow 1$$

DGKT lifts to 10d with smearred sources

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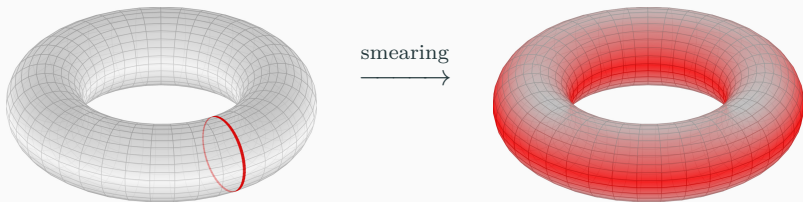


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Status of scale separation



Status of scale separation [DeWolfe, Giryavets, Kachru, Taylor '05]



Status of scale separation [Banks, van den Broek '07][McOrist, Sethi '12]

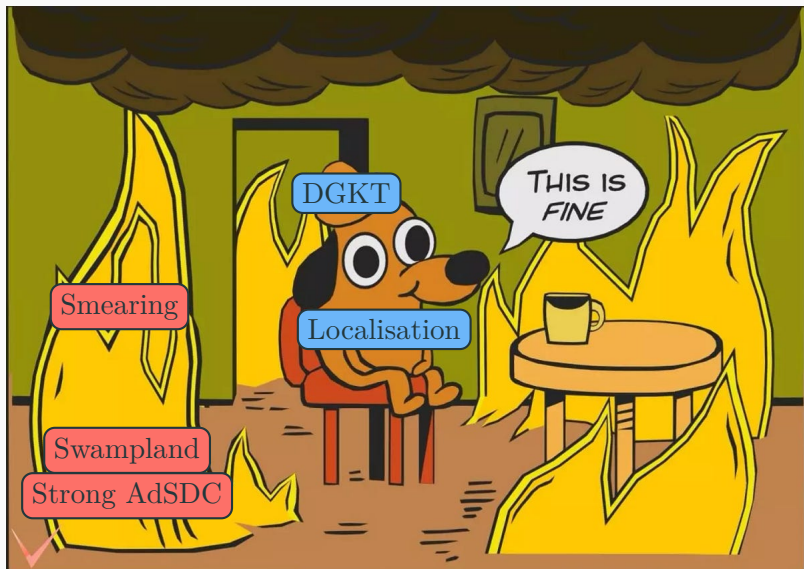


Status of scale separation [Junghans '20]

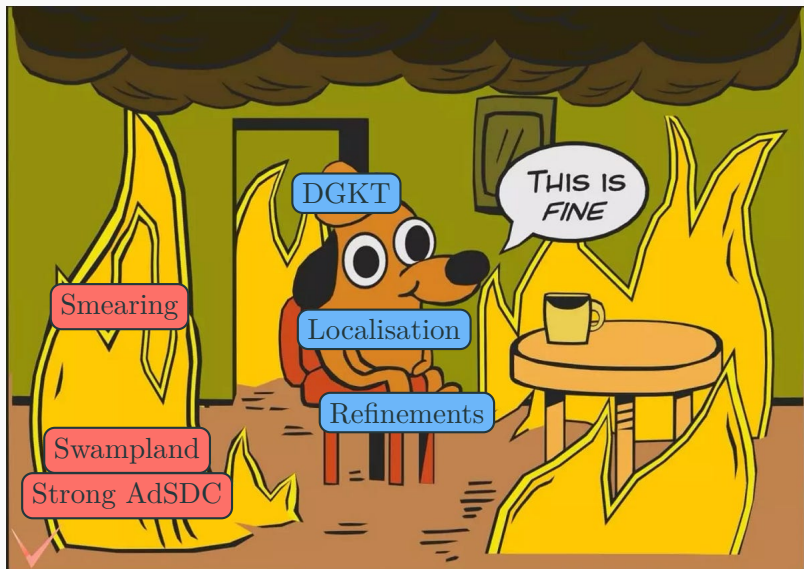
[Marchesano, Palti, Quirant, Tomasiello '20]



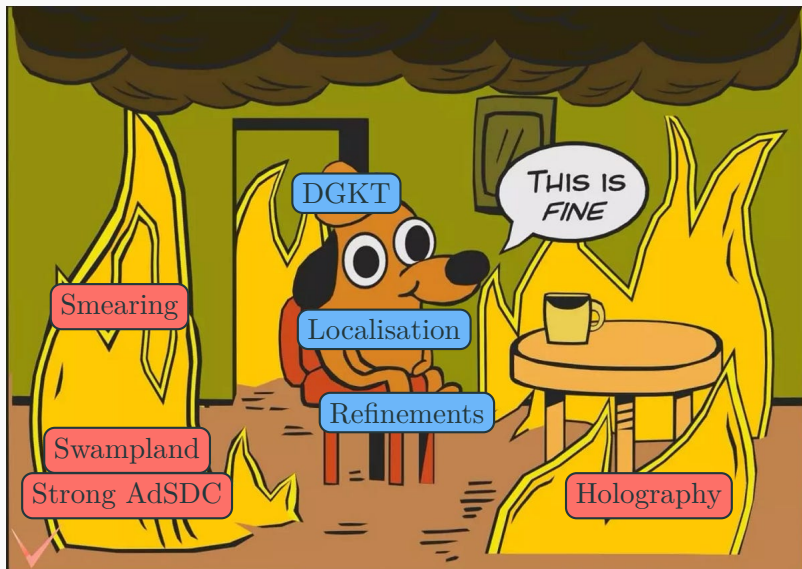
Status of scale separation [D. Lust, Palti, Vafa '19]



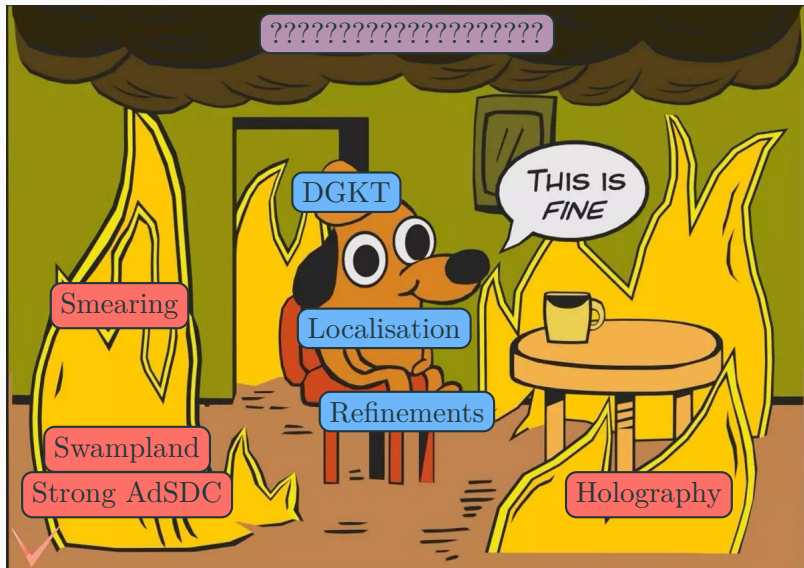
Status of scale separation [Buratti, Calderon, Mininno, Uranga '20]



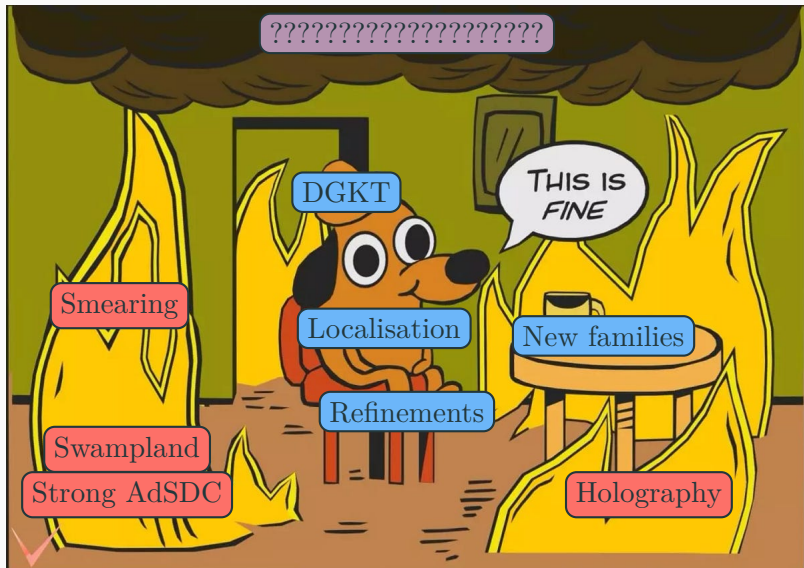
Status of scale separation [Polchinski, Silverstein '09] [Montero, Rocek, Vafa '23] [Collins, Jafferis, Vafa, Xu, Yau '22] [Apers, Conlon, Ning, Revello '22] ...



Status of scale separation



Status of scale separation



New families: Motivations

New families: Double T-duality [Banks, Van Den Broek '06]

- We consider $(T^2)^3/\Gamma$

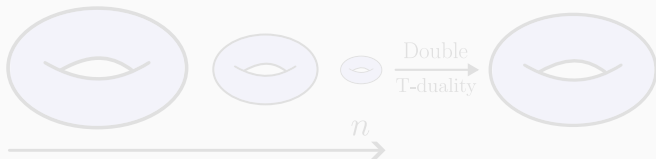
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- Scale separation: $r > 0, \quad r + s > 0$
- Weak coupling: $r + \frac{s}{2} > 0$

Can be satisfied even when $s < 0 \iff$ shrinking torus factor



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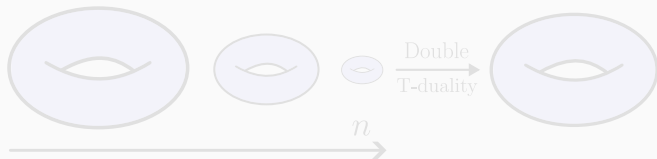
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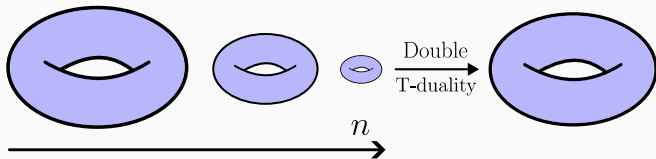
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Where T-duality brings us

- $K_K = -\log \left(i\kappa(T^1 - \bar{T}^1)(T^2 - \bar{T}^2)(T^3 - \bar{T}^3) \right), \quad \kappa = \mathcal{K}_{123}$
- $K_Q(\mathcal{P}^a)$

Double T-duality for $a = 1$: $T^1 \mapsto -\frac{1}{T^1}$

$$K_K \mapsto K_K + \log |T^1|^2$$

Kähler transf.: $K_K \mapsto K_K - F - \bar{F}, W \mapsto e^F W, \quad F = \log T^1$

$$W_{RR} = e_0 + e_a T^a + \frac{1}{2} \mathcal{K}_{abc} m^a T^b T^c + \frac{m}{6} \mathcal{K}_{abc} T^a T^b T^c$$



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Where T-duality brings us

$$\begin{array}{c} W_{\text{NS}} = h_{\mu} U^{\mu} \\ \downarrow \\ W_{\text{NS}} = h_{\mu} U^{\mu} T^1 \end{array}$$

- Quadratic term mixing U^{μ} and T^1

H-flux \mapsto Metric fluxes (rank-one)

Lost Romans mass but price to pay: Depart from CY

Back to 4d setup

$$W_{\text{RR}}, W_{\text{NS}} = h_\mu U^\mu + f_{a\mu} T^a U^\mu$$

$f_{a\mu} \in \mathbb{Z} \longleftrightarrow$ metric fluxes:

$$d\omega_a = -f_{a\mu} \beta^\mu, \quad d\alpha_\mu = -f_{a\mu} \tilde{\omega}^a$$

Tadpole: $[m^a f_{a\mu} + m h_\mu + m f_{a\mu} b^a] \beta^\mu + N_\alpha \delta_{\text{D6}}^\alpha - 4\delta_{\text{O6}} = 0$

$\implies e$ fluxes (G_4 and G_6) are not constrained
 \implies the m^a 's (G_2) not involved in $m^a f_{a\mu}$ either

Generalizes the toroidal construction of [Cribiori, Junghans, Van

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4d setup: Bilinear magic

$$V_F = \rho_A Z^{AB} \rho_B$$

[Bielleman, Ibañez, Valenzuela '15][Carta, Marchesano, Staessens, Zoccarato '16][Herraez, Ibañez, Marchesano, Zoccarato '18]

- Z^{AB} depends only on saxions
- The ρ_A 's are **flux-axion polynomials**
→ depend on fluxes and axions

Facilitates systematic search for vacua [Escobar, Marchesano, Staessens '18][Escobar, Marchesano, Staessens '19][Marchesano, Quirant '19][Marchesano, Prieto, Quirant, Shukla '20][Marchesano, Prieto, Wiesner '21]

New families: More details

4d setup: Bilinear magic

F-term ansatz [Marchesano, Prieto, Quirant, Shukla '20]:

$$\langle D_{T^a} W, D_{U^\mu} W \rangle_{\text{vac}} = \langle \lambda_K \partial_{T^a} K, \lambda_Q \partial_{U^\mu} K \rangle_{\text{vac}}$$

$$\rho_a = \ell_s^{-1} \mathcal{P} \partial_a K,$$

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Parameters \ Branch	\mathcal{P}	$\ell_s \rho_0$	\mathcal{Q}	$\ell_s \tilde{\rho} = m$	\mathcal{M}
SUSY	Free	$-\frac{3}{2} \mathcal{N}$	\mathcal{N}	$-10 \frac{\mathcal{P}}{\mathcal{K}}$	$-\frac{2}{3} \mathcal{P}$
non-SUSY	0	$-\frac{\mathcal{N}}{2} \left(1 - \frac{12}{S}\right)$	\mathcal{N}	0	$\frac{4\mathcal{P}}{S}$
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$$t^a \rho_{a\mu} = \ell_s^{-1} \mathcal{N} \partial_\mu K,$$

Parameters Branch	\mathcal{P}	$\ell_s \rho_0$	\mathcal{Q}	$\ell_s \tilde{\rho} = m$	\mathcal{M}
SUSY	Free	$-\frac{3}{2} \mathcal{N}$	\mathcal{N}	$-10 \frac{\mathcal{P}}{\mathcal{K}}$	$-\frac{2}{3} \mathcal{P}$
non-SUSY	0	$-\frac{\mathcal{N}}{2} \left(1 - \frac{12}{S}\right)$	\mathcal{N}	0	$\frac{4\mathcal{P}}{S}$
	$+\frac{\mathcal{N}}{2}$				
	$-\frac{\mathcal{N}}{2}$				

4d setup: Bilinear magic

F-term ansatz [Marchesano, Prieto, Quirant, Shukla '20]:

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Massless solutions

SUSY:	$Q = \mathcal{N},$	$l_s \rho_0 = -\frac{3}{2}\mathcal{N},$
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AdS vacua with $V|_{\text{vac}} = -12e^K Q^2$

10d interpretation:

- Half-flat $(\mathcal{W}_1, \mathcal{W}_2, \mathcal{W}_3)$ with $\mathcal{W}_3 = 0$
- Complements [Marchesano, Prieto, Quirant, Shukla '20] where the geometry was nearly-Kähler ($\mathcal{W}_1 \neq 0$)

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Vacuum equations

Metric flux Ansatz: $f_{a\mu} = \sigma_a \sigma_\mu \longrightarrow$ rank 1

\implies Closed expr. for stabilised axions and $Q(\text{fluxes})$, which sets vacuum energy

Saxions:

$$J_{ab} t^b|_{\text{vac}} = Q \left[4 \frac{\sigma_a}{\sigma_a t^a} - 3 \frac{\mathcal{K}_a}{\mathcal{K}} \right], \quad Q \partial_\mu K|_{\text{vac}} = \sigma_\mu \sigma_a t^a$$

$$\mathcal{K}_a \equiv \mathcal{K}_{abc} t^b t^c$$

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- To generate infinite families of sol.: Scaling symmetries

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- To generate infinite families of sol.: [Scaling symmetries](#)

Exact scaling symmetries

$$\text{If } \mathcal{K}_{abc} = \sigma_{(a}\eta_{bc)}, \quad \eta_{ab} = \eta_{ba}, \quad \sum_a \sigma_a \eta_{ab} = 0$$

$$Q \sim n^{2r}, \quad m^b \eta_{bc} \sim n^{r-s}, \quad m^a \sigma_a \sim \text{const.}, \quad \text{with } r \geq s \geq 0$$

$$\implies t^b \eta_{bc} \sim n^r, \quad t^a \sigma_a \sim n^s$$

$$\text{And scale separation } \underline{R_{\text{AdS}} M_{\text{KK}}} \sim n^{\frac{1}{2}(r-s)} \quad (g_s \sim n^{-r+\frac{3}{2}s})$$

$$\bullet X_6 = (T^2 \tilde{\times} X_4) / \Gamma$$

$$X_4 = T^4 \quad [\text{Cribiori, Junghans, Van Hemelryck, Van riet, Wrase '21}]$$

$$\text{or } X_4 = K3 \quad [\text{Caviezel, Koerber, Kors, Lurs, Tsimpis, Zagermann '09}]$$

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Ell. fibred: Approximate scaling symmetries

\mathcal{K}_{abc} more involved, scale sep. still possible?

$$\mathcal{K}_{Lab} = \eta_{ab}, \quad \mathcal{K}_{LLa} = \eta_{ab}c^b, \quad \mathcal{K}_{LLL} = \eta_{ab}c^a c^b, \quad \mathcal{K}_{abc} = 0$$

- Exact sym. if

$$m^a \sim n^{r-s}, \quad m^L \sim \text{const.}, \quad Q \sim n^{2r}, \quad c^a \sim n^{r-s}$$

- Hint for approx. scaling corrected by $\epsilon^a \equiv \frac{c^a}{m^a} \sim n^{s-r}$

→ Saxions as expansions in ϵ

$$t^L = t_{(0)}^L \left(1 + \Delta^L + \mathcal{O}(\epsilon^2) \right), \quad t^a = t_{(0)}^a - \frac{5}{3} \frac{t_{(0)}^L}{m^L} \left(\Delta^a + m^a \Delta^L + \mathcal{O}(\epsilon) \right)$$

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Perturbative stability

Compute Hessian ($\partial\partial V$) and diagonalize

→ very involved computation!!

But analytical expression is found!

→ Diagonal complex structure sector

Factorised cases:

$$\frac{m_s^2}{|m_{\text{BF}}|^2} = \left(8, \frac{280}{9}, 8, -\frac{8}{9}\right), \begin{cases} \text{SUSY:} & \frac{m_a^2}{|m_{\text{BF}}|^2} = \left(\frac{40}{9}, \frac{352}{9}, \frac{40}{9}, 0\right) \\ \text{non-SUSY:} & \frac{m_a^2}{|m_{\text{BF}}|^2} = \left(-\frac{8}{9}, \frac{160}{9}, \frac{160}{9}, 0\right) \end{cases}$$

Multiplicities $(1, 1, h_-^{1,1}(X_6), h^{2,1}(X_6))$

Universal DGKT spectrum [Marchesano, Quirant '19]

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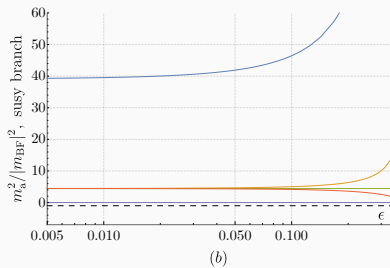
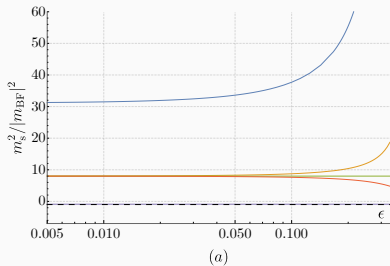
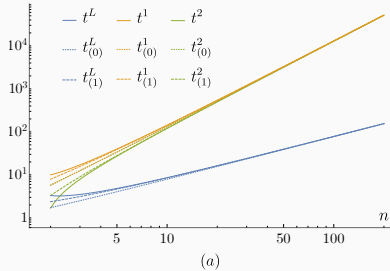
Universal DGKT spectrum [Marchesano, Quirant '19]

Simple elliptically fibered models:

→ T-duals of $\mathbb{P}_{(1,1,1,6,9)}$ ($h^{1,1} = 2$) and $\mathbb{P}_{(1,1,2,8,12)}$ ($h^{1,1} = 3$)

- **Confirm** the existence of infinite families
- **Check** analytics of approximate scaling symmetry
- **Compute** masses when not possible analytically
- **Crosscheck** of our analytical Hessian

Some numerics

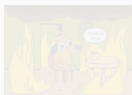


Conclusions

- **Scale separation** needed to justify 4d EFT



- Status **uncertain** in string theory



→ Progress on source localisation and past worries

→ Bottom-up constructions **clashes** with Swampland and holography

- We proposed new families of sol. in massless type IIA based on elliptically fibered CY with metric fluxes

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Outlooks

- **O-plane backreaction** [Junghans '20][Marchesano, Palti, Quirant, Tomasiello '20]
- Strong coupling regime to study M-theory uplift

$$R_{\text{AdS}} M_{\text{KK}} \sim n^{\frac{1}{2}(r-s)} \text{ and } g_s \sim n^{-r+\frac{3}{2}s}$$

So $r > s$ while $r < \frac{3}{2}s \implies$ **Scale sep.** and **strong coupling**

- In [Cribiori, Junghans, Van Hemelryck, Van Riet, Wrase '21] backreaction makes F_2 closed and suggest a **sourceless M-theory uplift geometry**

See also recently: [Van Hemelryck '24]

- Interesting to challenge conjecture in [Collins, Jafferis, Vafa, Xu, Yau '22]: “KK scale cannot be decoupled from internal curvature”
→ But sourceless scale sep. only possible if it is the case

[Gautason, Schillo, Van Riet, Williams '15]

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